

BH HORIZONS

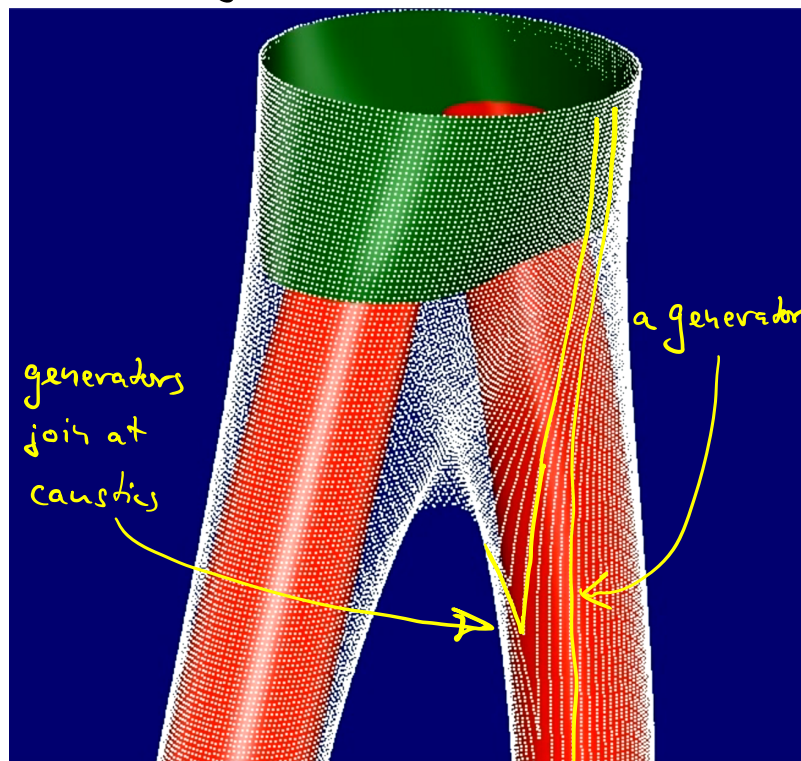
- Event horizons (global. Require 4-D data)
- Apparent horizons (local in time: Require 3-D data at $t = \text{const}$)

Event-Horizons

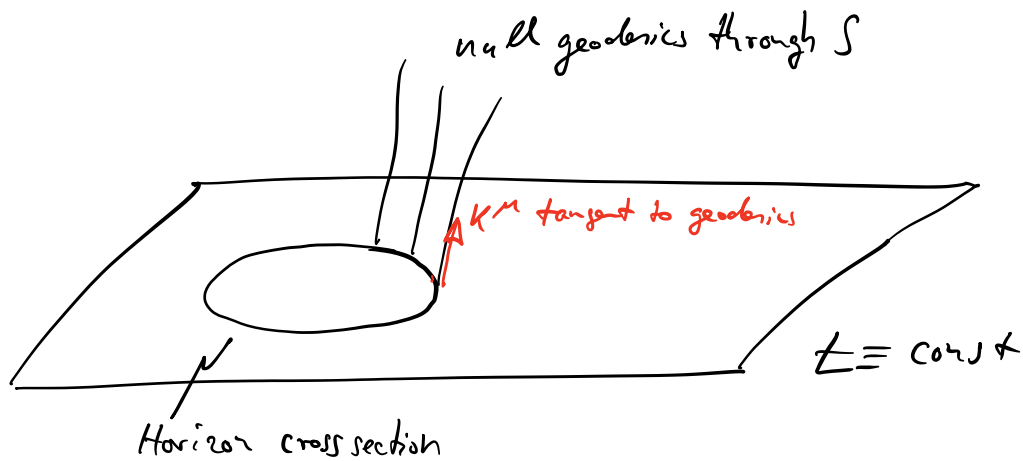
- foliated by null geodesics ("generators" of Horizon)
- which geodesics determined at late time
 - informally: those that don't fall into BH and neither escape
 - formally: boundary of past domain of dependence of $\mathcal{I}^+$
- new geodesics can enter horizon at caustics
- A_{EH} grows when geodesics diverge

Merger of two BHs "pair of pants"

space-time diagram



Key concept: expansion of null-geodesic congruence.



$$\Theta = \nabla_\mu k^\mu$$

Raychaudhuri's eqn for null congruences

$$\mathcal{L}_k \Theta = -\frac{1}{2} \Theta^2 - \underbrace{\sigma^{\mu\nu} \sigma_{\mu\nu}}_{\text{shear of congruence}} + \underbrace{\omega^{\mu\nu} \omega_{\mu\nu}}_{\text{twist of congruence}} - \underbrace{R_{\mu\nu} k^\mu k^\nu}_{\substack{=0 \text{ in vacuum} \\ \geq 0 \text{ for matter} \\ \text{satisfying} \\ \text{weak energy} \\ \text{condition}}}$$

$$\Rightarrow \frac{d\Theta}{d\lambda} \leq -\frac{1}{2} \Theta^2$$

↑ affine param
along geodesic

shear of
congruence
(= energy flux
across)

twist of congruence
= 0 b/c surface forming

$\Theta < 0 \Rightarrow \Theta \rightarrow -\infty$ in finite affine parameter

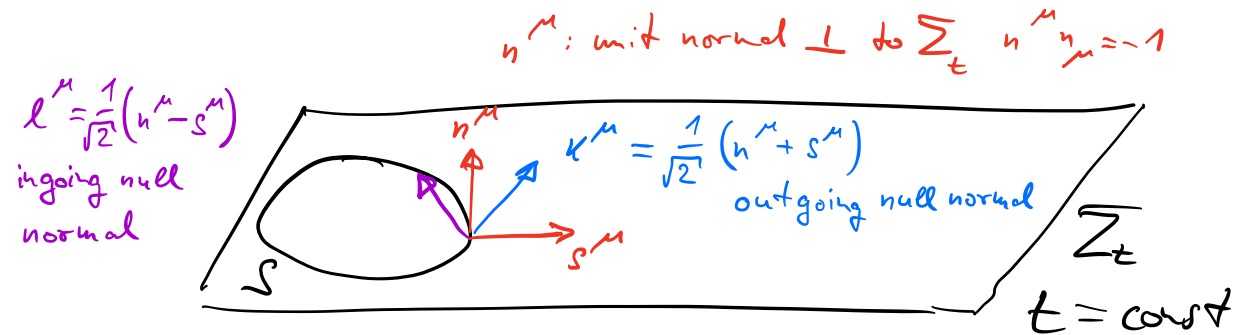
$\Rightarrow$ Penrose singularity then $\Rightarrow$ not on EH

$\Rightarrow \Theta \geq 0$ on EH

$\Theta = 0$ on EH $\cup$ constant area (non-growing)

Apparent Horizon

outermost marginally outer trapped surface MOTS



null:

$$\begin{aligned}
 k_\mu k^\mu &= \frac{1}{2} (h+s)(h+s) \\
 &= \frac{1}{2} \begin{pmatrix} h^2 & 2h \cdot s & s^2 \\ -1 & 0 & +1 \end{pmatrix} \\
 &= 0
 \end{aligned}$$

$$\text{MOTS: } \Theta_{(k)} = \nabla_\mu k^\mu = 0$$

In 3+1:

induced metric on Horizon:

$$m_{\mu\nu} = g_{\mu\nu} + k_\mu l_\nu + k_\nu l_\mu$$

$$\Theta_{(k)} = m^{\mu\nu} \nabla_\mu k_\nu$$

$$= \frac{1}{\sqrt{2}} m^{\mu\nu} \left(\nabla_\mu h_\nu + \nabla_\mu s_\nu \right)$$

$$= \frac{1}{\sqrt{2}} m^{\mu\nu} \left(-K_{\mu\nu} + \nabla_\mu s_\nu \right)$$

(*)

all tensors in (*) are spatial, so can switch to spatial indices.

Moreover $m_{ij} = \gamma_{ij} - s_i s_j$

$$\Rightarrow \Theta_{(k)} = \frac{1}{\sqrt{2}} \left(\mathcal{D}_i s^i - K + K_{ij} s^i s^j \right) \quad (**)$$

$\therefore$ to find AH (or more generally any MOTS), find topologically spherical surface S within Σ_t that satisfies $\Theta_{(k)} = 0$ everywhere on S .

AH (MOTS) depend only on data on Σ_t (γ_{ij}, K_{ij}).

For stationary BHs (where EH has $\Theta = 0$), $AH = EH$

For growing BHs (where EH has $\Theta > 0$), AH inside EH

Common AH
appears discontinuously

AHs of
individual BHs

